

# Paleocene latitude of the Kohistan–Ladakh arc indicates multistage India–Eurasia collision

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**We report paleomagnetic data showing that an intraoceanic Trans-Tethyan subduction zone existed south of the Eurasian continent and north of the Indian subcontinent until at least Paleocene time. This system was active between 66 and 62 Ma at a paleolatitude of  $8.1 \pm 5.6^\circ\text{N}$ , placing it 600–2,300 km south of the contemporaneous Eurasian margin. The first ophiolite obductions onto the northern Indian margin also occurred at this time, demonstrating that collision was a multistage process involving at least two subduction systems. Collisional events began with collision of India and the Trans-Tethyan subduction zone in Late Cretaceous to Early Paleocene time, followed by the collision of India (plus Trans-Tethyan ophiolites) with Eurasia in mid-Eocene time. These data constrain the total postcollisional convergence across the India–Eurasia convergent zone to 1,350–2,150 km and limit the north–south extent of northwestern Greater India to <900 km. These results have broad implications for how collisional processes may affect plate reconfigurations, global climate, and biodiversity.**

India | paleomagnetism | Neotethys | Himalaya | intraoceanic arc

Classically, the India–Eurasia collision has been considered to be a single-stage event that occurred at 50–55 million years ago (Ma) (1, 2). However, plate reconstructions show thousands of kilometers of separation between India and Eurasia at the inferred time of collision (3, 4). Accordingly, the northern extent of Greater India was thought to have protruded up to 2,000 km relative to present-day India (5, 6) (Fig. 1). Others have suggested that the India–Eurasia collision was a multistage process that involved an east–west trending Trans-Tethyan subduction zone (TTSZ) situated south of the Eurasian margin (7–9) (Fig. 1). Jagoutz et al. (9) concluded that collision between India and the TTSZ occurred at 50–55 Ma, and the final continental collision occurred between the TTSZ and Eurasia at 40 Ma (9, 10). This model reconciles the amount of convergence between India and Eurasia with the observed shortening across the India–Eurasia collision system with the addition of the Kshiroda oceanic plate. Additionally, the presence of two subduction systems can explain the rapid India–Eurasia convergence rates (up to  $16 \text{ mm a}^{-1}$ ) that existed between 135 and 50 Ma (9), as well as variations in global climate in the Cenozoic (11).

While the existence of the TTSZ in the Cretaceous is not disputed, the two conflicting collision models make distinct predictions about its paleolatitude in Late Cretaceous to Paleocene time; these can be tested using paleomagnetism. In the single-stage collision model, the TTSZ amalgamated with the Eurasian margin prior to ~80 Ma (12) at a latitude of  $\geq 20^\circ\text{N}$  (13, 14). In contrast, in the multistage model, the TTSZ remained near the equator at  $\leq 10^\circ\text{N}$ , significantly south of Eurasia, until collision with India (9) (Fig. 1).

No undisputed paleomagnetic constraints on the location of the TTSZ are available in the central Himalaya (15–17). West-erweel et al. (18) showed that the Burma Terrane, in the eastern Himalaya, was part of the TTSZ and was located near the equator at ~95 Ma, but they do not constrain the location of the

TTSZ in the time period between 50 and 80 Ma, which is required to test the two collision hypotheses. In the western Himalaya, India and Eurasia are separated by the Bela, Khost, and Muslimbagh ophiolites and the 60,000 km<sup>2</sup> intraoceanic Kohistan Ladakh arc (19, 20) (Fig. 1). These were obducted onto India in the Late Cretaceous to Early Paleocene (19), prior to the closure of the Eocene to Oligocene Katawaz sedimentary basin (20) (Fig. 1). The Kohistan–Ladakh arc contacts the Eurasian Karakoram terrane in the north along the Shyok suture and the Indian plate in the south along the Indus suture (21) (Fig. 1). Previous paleomagnetic studies suggest that the Kohistan–Ladakh arc formed as part of the TTSZ near the equator in the early Cretaceous but provide no information on its location after 80 Ma (22–25). While pioneering, these studies lack robust age constraints, do not appropriately average paleosecular variation of the geodynamo, and do not demonstrate that the measured magnetizations have not been reset during a subsequent metamorphic episode.

## Paleomagnetism of the Khardung Volcanics

Paleocene volcanic rocks from the Kohistan–Ladakh arc provide an unrivaled opportunity to use paleomagnetism to reconstruct the paleolatitude of the TTSZ shortly before onset of collision and test the two conflicting models. We sought to determine the paleolatitude of the Paleocene Khardung volcanics on the northern margin of the Kohistan–Ladakh arc in Ladakh, India. The ~3,000-m-thick stratigraphy of the Khardung volcanics formed between 70 and 60 Ma (26, 27) and comprises rhyolitic and andesitic

## Significance

**We present paleomagnetic constraints on the latitude of an intraoceanic subduction system that is now sutured between India and Eurasia in the western Himalaya. Our results demonstrate that the India–Eurasia collision was a multistage process involving at least two subduction systems rather than a single-stage event. This resolves the discrepancy between the amount of convergence and the observed crustal shortening in the India–Eurasia collision system, as well as the 10–15 Ma time lag between collision onset in India and the initiation of collision-related deformation and metamorphism in Eurasia. The presence of an additional subduction system in the Neotethys ocean explains the rapid India–Eurasia convergence rates in the Cretaceous and global climate variations in the Cenozoic.**

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The authors declare no competing interest.

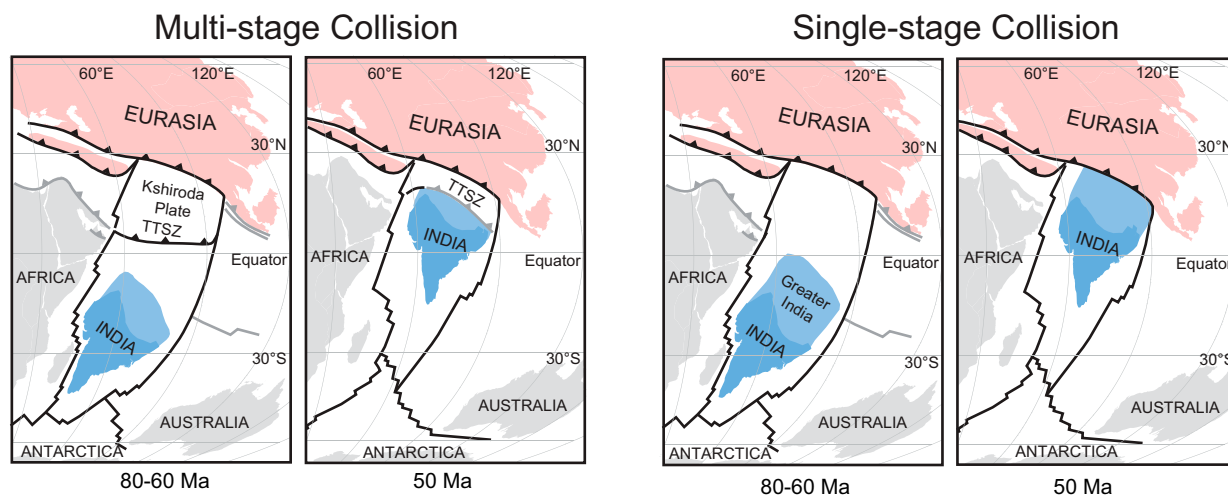
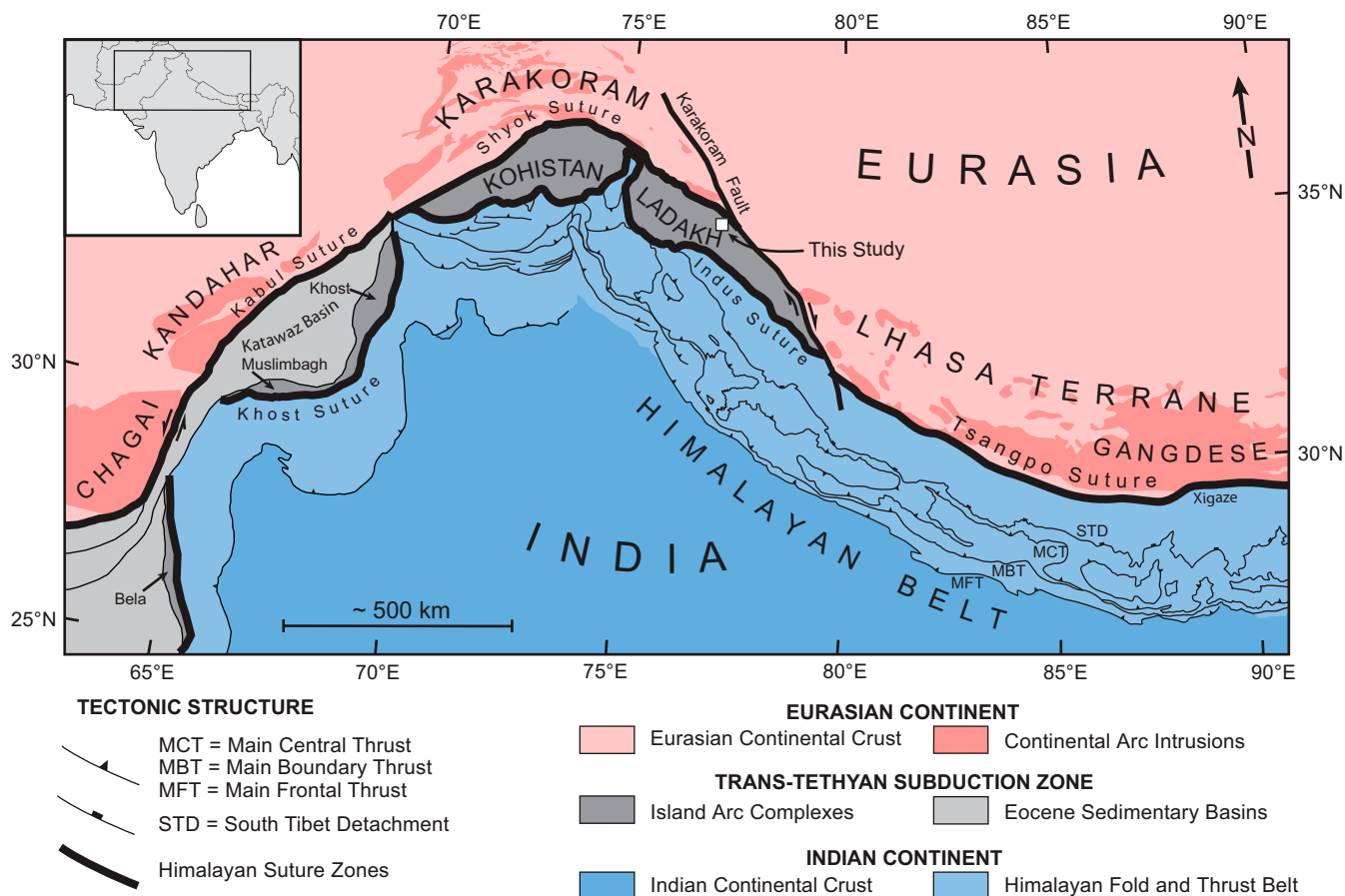
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**Fig. 1.** The first panel is an overview map of tectonic structure of the Karakoram–Himalaya–Tibet orogenic system. Blue represents India, red represents Eurasia, and the Kohistan–Ladakh arc (KLA) is shown in gray. The different shades of blue highlight the deformed margin of the Indian plate that has been uplifted to form the Himalayan belt, and the zones of darker red within the Eurasian plate highlight the Eurasian continental arc batholith. Thick black lines denote the suture zones which separate Indian and Eurasian terranes. The tectonic summary panels illustrate the two conflicting collision models and their differing predictions of the location of the Kohistan–Ladakh arc. India is shown in blue, Eurasia is shown in red, and the other nearby continents are shown in gray. Active plate boundaries are shown with black lines, and recently extinct boundaries are shown with gray lines. Subduction zones are shown with triangular tick marks.

lava flows, tuffs, ignimbrites, agglomerate, and minor clastic sediments at the top of the section (*SI Appendix, Figs. S1 and S2*).

Zircons from ash layers and lava flows distributed throughout the upper 1,000 m of the exposed section were dated using U–Pb chemical abrasion–isotope dilution–thermal ionization mass spectrometry (CA-ID-TIMS) geochronology (*SI Appendix, Table*

*S1 and Fig. S4*). The U–Pb ages are consistently younger from the bottom to the top of the stratigraphy indicating that there are no major repetitions or duplications caused by unidentified fault zones. Zircons separated from a rhyolite flow (LB13–17) sampled from the bottom part of the studied section yield a Th-corrected  $^{206}\text{Pb}/^{238}\text{U}$  weighted mean eruption/deposition age of  $65.038 \pm 0.12$

Ma ( $2\sigma$  external uncertainty), and zircons from the intermediate ash layer (KA1) yield a Th-corrected  $^{206}\text{Pb}/^{238}\text{U}$  weighted mean eruption/deposition age of  $62.097 \pm 0.079$  Ma. Two samples yielded inherited zircon populations, such that only a maximum depositional age could be constrained from the youngest zircon in the sample. The rhyolite flow sampled from the top of the section (LB13–16) has a Th-corrected  $^{206}\text{Pb}/^{238}\text{U}$  zircon maximum depositional age of  $61.636 \pm 0.11$  Ma, and the interlayered ash horizon (KA4D) at the bottom of the sampled stratigraphy has a Th-corrected  $^{206}\text{Pb}/^{238}\text{U}$  zircon maximum depositional age of  $66.100 \pm 0.085$  Ma. Based on these results the absolute time within our sampled section is 61.64–66.10 Ma, a time span of 4.46 Ma (Fig. 2).

We obtained a total of 191 oriented 2.2-cm-diameter cores from 21 sites across the dated section. At each site, the orientations of multiple bedding planes were measured, and a site-specific mean was calculated and applied to the magnetic data to correct for bedding tilt (Fig. 3). The sites used to estimate the paleolatitude each consist of 5–13 core samples from single flows that should have acquired a near-instantaneous (a few hours) thermoremanent magnetization from the local magnetic field as they cooled. For most samples, low-coercivity/low-temperature (LCT) overprints were removed using alternating field (AF) demagnetization steps between 1 and 20 mT and thermal demagnetization steps between 100 °C and 300 °C (*SI Appendix, Figs. S7 and S8*). Stable high-temperature magnetization components consistent with magnetite were isolated in the thermal demagnetization steps between 500 °C and 580 °C. In a small subset of the samples, stable, high-temperature components unblocking between 600 °C and 680 °C consistent with hematite were also present, oriented parallel to the magnetite component (*SI Appendix, Fig. S7 and Dataset S1*).

We confirmed the absence of significant posteruption remagnetization using a reversal test and two conglomerate tests consisting of 69 cores from clasts from two intraformational conglomerate units near the top and bottom of the sampled section (29) (*SI Appendix, Figs. S9 and S10*). The high-temperature magnetization directions from the clasts in the conglomerate units are random with  $\geq 95\%$  certainty (30), therefore indicating a lack of total remagnetization since deposition. The paleomagnetic measurements of the bedded volcanics define two antipodal populations (Fig. 3) extending over three geomagnetic polarity reversals that correlate to chrons C29n, C28r, C28n, C27r, and C26r, showing that they have not been overprinted (28) (Fig. 2). Our successful conglomerate tests and reversal test unequivocally demonstrate that the HT1 and HT2 directions in the samples are primary (29, 31).

We obtained a Fisher mean paleomagnetic pole located at a latitude of  $64.0^\circ\text{N}$  and a longitude of  $266.4^\circ\text{E}$  with a 95% confidence angle of  $A_{95} = 5.6^\circ$  (*SI Appendix, Fig. S12*). Secular variation should have been successfully averaged by our measurements given the  $>4$ -Ma time span recorded within the section (including three reversals) and the fact that the value of  $A_{95}$  is within the observed range for the modern geomagnetic field ( $5.3^\circ < A_{95} < 13.3^\circ$ , for  $n = 18$ ) (29, 31, 32). The measured paleomagnetic pole constrains the paleolatitude of the Kohistan–Ladakh arc to  $8.1 \pm 5.6^\circ\text{N}$  at 61.64–66.10 Ma, significantly south of the Eurasian margin which was situated at  $21.2 \pm 2.1^\circ\text{N}$  at the same time (14). This constrains the location of the TTSZ to 600–2,300 km south of the Eurasian margin in the Paleocene (Fig. 4).

## Discussion

Our results indicate that until at least  $\sim 61.6$  Ma, and probably until 50–55 Ma, the Neotethys ocean was subducted along two separate subduction systems, an active continental system at the southern Eurasian margin and the TTSZ located in the equatorial Neotethys (7–9) (Fig. 4). These findings support multistage tectonic evolution models for the western Himalaya in which the India–TTSZ–Eurasia collision began with collision between the

TTSZ and the northern margin of India in the Late Cretaceous to Paleocene and ended with final continental collision in the mid-Eocene (9, 10) (Fig. 4).

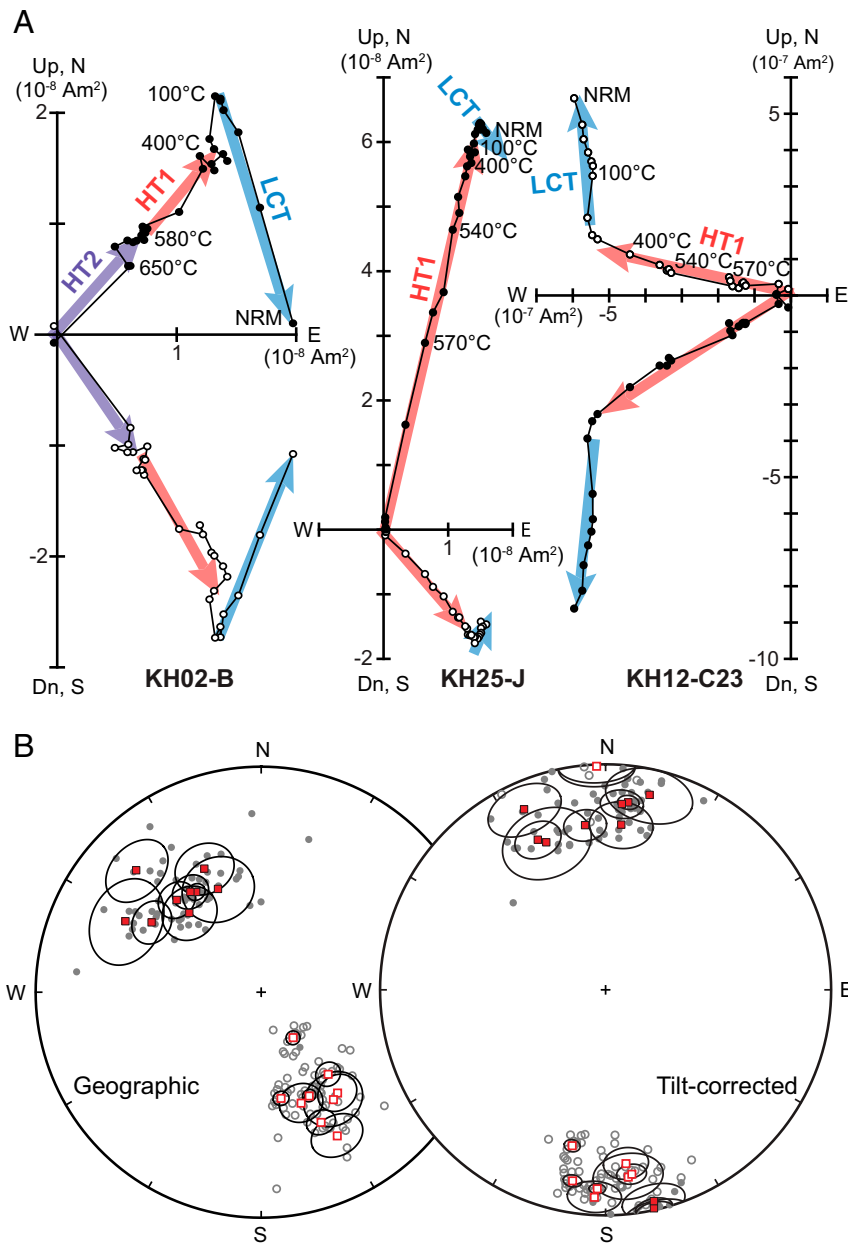
The termination of intraoceanic subduction during the first stage of collision resulted in the Late Cretaceous/Early Paleocene obduction of the Bela, Muslimbagh, and Khost ophiolites onto the northwestern margin of India (19), as well as the formation of the 50–55 Ma Indus suture zone between the Kohistan–Ladakh arc and India (10). It also caused a reduction in the rate of northward motion of India at  $52 \pm 4$  Ma (9). The final India–TTSZ–Eurasia continent–continent collision occurred along the Shyok–Tsangpo suture, not the Indus–Tsangpo, and the age of this final collision is constrained to  $40.4 \pm 1.3$  Ma by geochemical and isotopic changes in the Kohistan–Ladakh batholith (10). This explains why the India–Eurasia convergence rate continued to decrease until 40–45 Ma (33).

Recent Cretaceous age estimates for the Kohistan–Ladakh arc–Eurasia collision based on sediment provenance interpretations of detrital zircon U–Pb age data (34, 35) are incompatible with our results. We note that Borneman et al. (35) report that Eurasia-derived detrital zircons were deposited on Kohistan–Ladakh arc rocks at 80 Ma despite previous work showing that the sediments they sampled were deposited on the Eurasian margin, not the Kohistan–Ladakh arc (36–38). Similarly, Najman et al. (34) interpret a uniquely Eurasian provenance for seven 200–220 Ma grains from Paleocene Indian passive margin sediments, but these could have been reworked from Late Triassic and Early Jurassic Indian passive margin sediments which contain similar age populations (39, 40). Unlike our paleomagnetic data, these detrital zircon provenance investigations are limited by the fact that the vast area of Greater Indian landmass, now lost beneath Tibet, likely contributed detritus of unknown age to passive margin sediments before the collision.

Geological evidence for a multistage collision history of the Himalaya is best preserved in the western Himalaya, where the Kohistan–Ladakh arc and other ophiolite bodies clearly demarcate the TTSZ, and east of the Namche Barwa syntaxis, where the TTSZ is represented by the Burma Terrane (18). The exposure of the Kohistan–Ladakh arc is  $\sim 200$  km wide in northeast Pakistan but reduces to  $< 20$  km in southwestern Tibet where it disappears at the intersection of the great counter thrust system and the Karakoram fault (41) (Fig. 1). In south central Tibet the record of the TTSZ is fragmentary due to the large-scale back-thrusting along the great counter thrust system that obscures the complex evolution of the Tsangpo suture zone (42–44). It is not surprising that the geological record of the multiple stages of the India–TTSZ–Eurasia collision is limited in the central part of the Himalayan belt, where shortening and underthrusting were presumably highest.

The dismembered ophiolites in the Tsangpo suture zone could have formed in the TTSZ (7, 45), or in the forearc of the continental margin (17, 46). While the Xigaze ophiolite likely formed on the southern edge of Eurasia, it has recently been suggested that it moved southward in the Cretaceous during back arc extension after 85–90 Ma, becoming part of the TTSZ (44). However, subduction along the TTSZ initiated in the Jurassic [ $\sim 154$  Ma (47)] and not in the Cretaceous. Alternatively, if the TTSZ and Xigaze units are unrelated, fragments of the TTSZ could remain unidentified in the accretionary mélange south of the Xigaze ophiolite or elsewhere in the less well studied areas of the suture zone. Regardless, our results require that a major oceanic basin existed in between India and Eurasia at the same time Cretaceous arc detritus was deposited in Indian passive margin sediments in southeastern Tibet at 58.5–60 Ma (48), making the TTSZ the likely sedimentary source rather than Eurasia.

The near-equatorial location of the TTSZ in the Paleocene implies that a significant proportion of the 2,800–3,600 km India–Eurasia convergence since 50–55 Ma (49) can be accounted



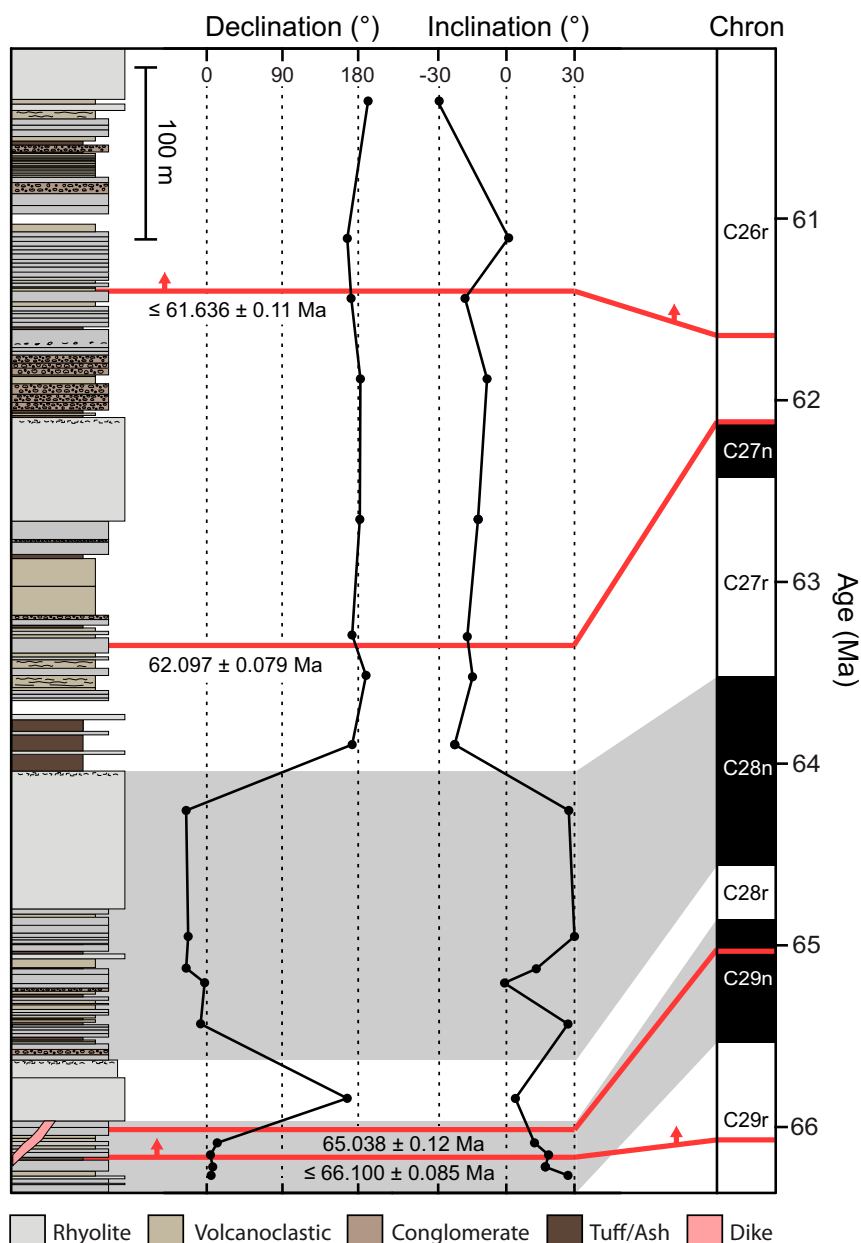
**Fig. 2.** (A) Orthographic projection diagrams showing AF and thermal demagnetization of three representative samples KH02-B, KH25-J, and conglomerate clast KH12-C23. Data are presented in geographic coordinates; closed symbols represent north-south-east-west projections, and open symbols represent up-down-east-west projections. Interpretations are shown with colored arrows: each arrow reflects a direction vector corresponding to components inferred from PCA. LCT = low-coercivity/low-temperature overprint, HT1 = high-temperature magnetite, HT2 = high-temperature hematite. (B) Stereographic equal-area projections showing HT1 directions for each sample (gray circles) and their site-means (red squares) with associated 95% confidence angles (black ellipses). Data are presented in geographic coordinates (*Left*) and after tilt correction (*Right*). Upward directions are denoted with open symbols, and downward directions are denoted with filled symbols.

for by precollisional subduction of the Kshiroda oceanic plate beneath Eurasia until  $\sim 40$  Ma rather than by thickening and extrusion of Indian and Eurasian continental crust. Indeed, the extrusion of southeast Asia started later, at  $\sim 36$  Ma (50), and metamorphism and melting in the High Himalaya began after 40 Ma (51). Comparison of our data to the well-constrained paleolatitudes of the Eurasian margin and India (13, 14) suggests that the Kshiroda plate was  $1450 \pm 850$  km wide, at the time of the TTSZ-India collision (Fig. 4). Therefore, the convergence accommodated by deformation in the Himalayan orogen was 1,350–2,150 km, similar to the 1,050–1,950 km combined Eurasian

and Indian shortening observed across western and eastern Tibet and the Himalayan fold and thrust belt (6, 52). Our results also constrain the size of Greater India in the western Himalaya to  $<900$  km, consistent with the observed  $\sim 600$ -km extent of Indian continental lithosphere underthrust beneath Eurasia (53) and reconstructions of Greater India based on Cretaceous paleomagnetic data (14).

In conclusion, we present robust and reliable paleomagnetic data constraining the location of the TTSZ to  $8.1 \pm 5.6^\circ \text{N}$  in the Paleocene, 600–2,300 km south of the Eurasian margin. Our results require that two subduction zones were active throughout





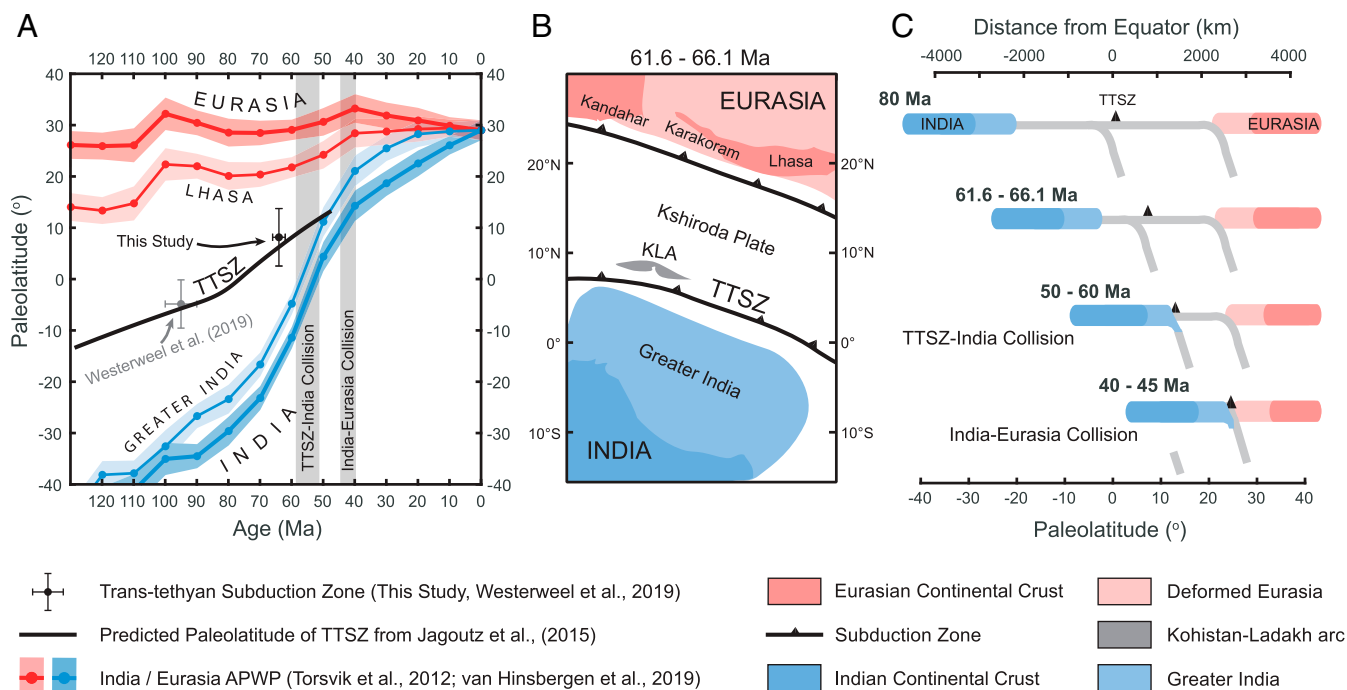
**Fig. 3.** Stratigraphic column of upper ~1,000 m of Khardung volcanics where we collected paleomagnetic and geochronology samples. Rock types are denoted by the color of the blocks (gray = rhyolite, yellow = volcanoclastic, light brown = conglomerate, dark brown = tuff/ash, red = intermediate dike), the horizontal extent of the blocks represents the relative erosive prominence of the units in the field, and breaks in the section reflect small areas with no exposure. The site-mean declinations and inclinations are plotted against stratigraphic height, and our four U-Pb ages ( $2\sigma$  external uncertainty) are shown as red lines; up arrows indicate maximum depositional ages. The gray shaded regions show the correlation of the magnetic reversals in the sequence to the documented C29n, C28r, C28n, C27r, and C26r chrons which have been plotted with their age in the right-hand column (28).

the closure of the Neotethys until the Paleocene. We have also shown that the India–Eurasia collision was a multistage process that began with the accretion of the TTSZ onto India in the Late Cretaceous to Paleocene and ended with continent–continent collision in the Eocene at ~40 Ma. The north–south extent of Greater India in the west was ~900 km, and the final collision occurred along the Shyok–Tsangpo suture zone, not the Indus–Tsangpo. The combined activity of both Neotethyan subduction systems explains the anomalously rapid motion of India in the Late Cretaceous (9), and the low-latitude obduction of ophiolites associated with the two stages of India–Eurasia collision caused the global cooling observed throughout the Cenozoic (11). Our study demonstrates that the peripheries of the Himalayan belt

provide crucial insight into the geological evolution of the India–TTSZ–Eurasia collision system that is difficult to discern in the central part of the orogen.

### Materials and Methods

**U-Pb Zircon Geochronology.** Zircons were separated from each sample using standard crushing and density separation techniques. U-Pb geochronology was completed using the CA-ID-TIMS technique at Massachusetts Institute of Technology (MIT), following methods slightly modified from Mattinson (54) and outlined in the appendix to Eddy et al. (55). Zircons were first annealed at 900 °C and 1 atm for 60 h. Subsequently, individual zircons were loaded into Teflon microcapsules with 100–125  $\mu$ L of 29M HF. The microcapsules were then loaded into a Parr dissolution vessel and held at 215 °C for 12–13 h.



**Fig. 4.** (A) Paleolatitude of the TTSZ constrained at the KLA (this study) and the Burma Terrane (18) compared to paleomagnetic plate reconstructions of Indian and Eurasian terranes (4, 14) and the predicted location of the TTSZ throughout the closure of the Neotethys ocean. (B) Paleogeographic map of the location of Kohistan-Ladakh arc relative to India and Eurasia in the Paleocene. The position of the Kohistan-Ladakh arc is reconstructed using our paleomagnetic pole from the Khardung volcanics, and the locations of Indian and Eurasian tectonic blocks are from the plate reconstruction of van Hinsbergen et al. (14). (C) Cross-section illustration showing the plate tectonic configuration of the India-TTSZ-Eurasia collision at 80 Ma, 61.6–66.1 Ma (constrained by our data), 50–60 Ma, and 40–45 Ma.

The resulting solutions were discarded, and each individual zircon was repeatedly rinsed in  $H_2O$ , and 6N HCl. After rinsing, approximately  $\sim 0.01$  g of EARTHTIME  $^{202}Pb$ - $^{205}Pb$ - $^{233}U$ - $^{235}U$  isotopic tracer (56) and 75–100  $\mu L$  of 29M HF were added to each microcapsule. The microcapsules were then reloaded into a Parr dissolution vessel and held at 215  $^{\circ}C$  for 48–60 h for total digestion. The solutions were subsequently dried down and dissolved in 6N HCl at 180  $^{\circ}C$  for  $\sim 12$  h to convert the samples to chloride form. Uranium and Pb were purified from the dissolved sample with AG-1 X8 200–400 mesh anion exchange resin using methods modified from Krogh et al. (57). Samples were first loaded onto 50  $\mu L$  anion exchange columns in 50–75  $\mu L$  of 3N HCl and rinsed dropwise to remove trace elements. Then Pb and U were eluted using 200  $\mu L$  of 6N HCl and 250  $\mu L$  of  $H_2O$ , respectively. Samples were dried down with a microdrop of 0.05M  $H_3PO_4$  prior to analysis via TIMS.

All of the data presented in this study were analyzed on the IsotopX Phoenix TIMS or the VG Sector 54 TIMS at MIT. Lead was run as a metal and measured by peak-hopping on a Daly photomultiplier. Uranium was analyzed as  $UO_2$  and was measured statically on a series of Faraday cups. Measured ratios were corrected assuming an  $^{18}O/^{16}O$  of  $0.00205 \pm 0.00004$  (2  $\sigma$ ), corresponding to the modern atmospheric value of Nier (58). Corrections for mass-dependent fractionation of U were done using the known ratio of  $^{233}U/^{235}U$  in the ET535 isotopic tracer and assuming a  $^{238}U/^{235}U$  of  $137.818 \pm 0.045$  (2  $\sigma$ ), which represents the mean value of  $^{238}U/^{235}U$  measured in natural zircon (59). Corrections for Pb fractionation were done using an  $\alpha$  (%) calculated from repeat runs of the NBS 981 Pb isotopic standard for the IsotopX Phoenix TIMS and an  $\alpha$  calculated from 53 Pb measurements on the Sector 54 TIMS of the ET2535 isotopic tracer, which contains a known  $^{202}Pb/^{205}Pb$  ratio.

A well-known problem in the measurement of small amounts of Pb by TIMS is the effect of isobaric interferences. Known isobaric interferences include  $BaPO_4$  and  $Tl$  and were corrected by measuring masses 201 and 203, assuming that they represent  $^{201}BaPO_4$  and  $^{203}Tl$ , and using the natural abundances of  $^{202}BaPO_4$ ,  $^{204}BaPO_4$ ,  $^{205}BaPO_4$ , and  $^{205}Tl$  to correct the measurements of masses 202, 204, and 205. These corrections were often minor and have no effect on our data interpretations.

A correction for common Pb ( $Pb_c$ ) was done by assuming that all  $Pb_c$  is from laboratory contamination and using the measured  $^{204}Pb$  and a laboratory  $Pb_c$

isotopic composition to subtract the appropriate mass of  $Pb_c$  from each analysis. We consider the assumption that all measured  $Pb_c$  is from laboratory contamination to be robust because the typical  $Pb_c$  seen in zircon analyses ( $< 1$  pg) is comparable to the mass of  $Pb_c$  seen in procedural blanks. One hundred forty-nine procedural blanks were used to quantify the  $Pb_c$  isotopic composition at MIT of  $^{206}Pb/^{204}Pb = 18.13 \pm 0.96$  (2  $\sigma$ ),  $^{207}Pb/^{204}Pb = 15.28 \pm 0.60$  (2  $\sigma$ ),  $^{208}Pb/^{204}Pb = 37.04 \pm 1.77$  (2  $\sigma$ ).

A correction for initial secular disequilibrium in the  $^{238}U$ - $^{206}Pb$  system due to the exclusion of Th during zircon crystallization (e.g., 60) was made for each analysis using a ratio of zircon/melt partition coefficients ( $f_{Th}$ ) of 0.119. This value was determined from coexisting zircon rims/surfaces and high- $SiO_2$  glass from a dacitic lava from Mt. St. Helens (61). We view this as the best available analog for silicic, hydrous, arc magmatism. Nevertheless, we have applied a generous uncertainty of  $\pm 1$  (2  $\sigma$ ) for the calculated [Th/U] magma.

All data reduction was done with the Tripoli and ET\_Redux software packages (62) using the algorithms presented by McLean et al. (63). The U decay constants are from Jaffey et al. (64). All isotopic data are presented in *SI Appendix, Table S1*. Our age interpretations for each sample include weighted mean eruption/deposition ages for those samples that contain a coherent age population (LB13–17 and KA1) and maximum depositional ages for samples that did not include a coherent population (KA4D and LB13–16). These samples may have incorporated xenocrysts during eruption/deposition, and we conservatively use the youngest grain as a maximum age. All uncertainties within the text are reported as external 2  $\sigma$  to aid in comparison with other geochronologic datasets produced by other techniques within the orogenic belt.

**Paleomagnetism.** Core samples were drilled in the field using a water-cooled electric hand drill. Cores were oriented in the field using an ASC Industries Pomeroy orienting fixture and extracted from the outcrop using nonmagnetic brass tools. One paleomagnetic specimen was cut from each core sample at MIT using an ASC Scientific dual-blade rock saw. Measurements of the natural remanent magnetization were obtained in the MIT Paleomagnetism Laboratory using a 2G Enterprises Superconducting Rock Magnetometer equipped with an automated sample handler (65) inside a mu-metal magnetically shielded room with  $< 200$ -nT DC field. The specimens were subjected to step-wise AF and thermal demagnetization. AF steps were applied in increments of

4 mT up to 20 mT and followed by thermal steps starting at 100 °C and increasing up to 600 °C or 680 °C in variable interval sizes from a maximum of 100 °C and minimum of 5 °C close to the Curie temperatures of the suspected principal magnetic carriers, magnetite (580 °C) and hematite (680 °C). Stable components of magnetization were isolated using principal component analysis (PCA) (66) (see *SI Appendix, Fig. S6* for examples of typical sample demagnetizations and *Dataset S1* for all fitting data). The domain state of magnetite was assessed using hysteresis curves (*SI Appendix, Figs. S4 and S5*) that were measured for a representative subset of bedded volcanic samples (KH01B, KH03F, KH14A, and KH25A) on an ADE model 1660 vibrating sample magnetometer (VSM) in the Ross Laboratory, MIT Department of Materials Sciences and Engineering.

Site-mean magnetization directions were calculated using Fisher (67) statistics and corrected for bedding tilt (Fig. 2 and *SI Appendix, Table S3*). We calculated Northern Hemisphere virtual geomagnetic pole (VGP) positions for each site (*SI Appendix, Table S3 and Fig. S11*), then we took the Fisher (67) mean of the VGP data to obtain the overall paleomagnetic mean pole and associated  $A_{95}$  error envelope. We used the quantile–quantile method (68) to show that the VGP distribution is consistent with a Fisher model (*SI Appendix, Fig. S10*).

To test for postdepositional remagnetization, we performed two conglomerate tests and a reversal test (29). The conglomerate test samples consisted of cores from 46 clasts from an intraformational conglomerate unit at site KH12 (near the top of the section) and 23 clasts from another intraformational conglomerate unit at site KP3 (near the bottom of the section) (*SI Appendix, Fig. S8*). We performed a Watson (30) test for randomness on the

distribution of tilt-corrected magnetic components from the clast specimens in each conglomerate unit. Both sets of conglomerate clast magnetization directions passed the conglomerate test, indicating that the clast magnetization directions were primary (see *SI Appendix* for further details). We used the bootstrap reversal test of Tauxe (69) to demonstrate that the two populations of site-mean magnetization directions from the volcanic sequence are antipodal to each other, therefore passing the reversal test (*SI Appendix, Fig. S9*). All paleomagnetic data reduction and interpretation were carried out using the PmagPy software package (70).

**Data Availability.** All study data are included in the article and supporting information.

**Note Added in Proof.** The description of the geochronology methods in this paper is similar to that in ref. 71.

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Supplementary Information for:

## **Paleocene latitude of the Kohistan-Ladakh arc indicates multi-stage India-Eurasia collision**

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**This PDF file includes:**

- Supplementary Text
- Supplementary Figures S1 – S12
- Supplementary Tables S1 – S3
- References

## **Geology of the Khardung volcanics:**

The Khardung volcanics consist of rhyolitic lava flows, welded tuffs, ash and volcano-sedimentary conglomerates locally intruded by meter-scale dykes of intermediate composition. The volcanic stratigraphy is deposited onto and locally intruded by the plutonic rocks of the Ladakh batholith, and contacts metasedimentary and metavolcanic rocks of the Shyok suture zone along the Khalsar fault (See Supplementary Figure S1). Rhyolite flows are 1 – 5 m thick and are typically very fine grained, often with chilled margins at their base and upper surface (See Supplementary Figure S2). Some rhyolites exhibit vesicular or porphyritic textures but most are equigranular and massive. Four 30 – 60 m thick rhyolite units are exposed in the sampled section and have a slightly coarser grain size and no discernable internal flow or bedding structures. Throughout the section, volcano-sedimentary deposits are interbedded with the rhyolites becoming more frequent in the uppermost 300 m. These laterally extensive deposits are typically 30 – 50 cm thick with angular, poorly sorted, lithic clasts, < 0.5 – 5 mm in diameter. The sediments provide clear bedding planes to determine the paleohorizontal plane and often exhibit graded bedding, trough-crossbedding and localized channelization indicating that the sequence has not been overturned. The conglomerates sampled for our conglomerate tests are both ~ 4 m thick and contain sub-angular intermediate-rhyolitic clasts, 10 – 80 cm diameter, and supported in a matrix of fine-grained lithic sand and ash particles (See Supplementary Figure S2). Ash-flow tuffs, typically 3 – 8 m thick with 5 – 15 cm volcanic bombs supported in very fine and glassy matrix, are also present in the uppermost 300 m of the section. The metamorphic overprint in the sediments is weak to absent with only minor deuteritic metamorphism related to the dewatering and degassing of the volcanic flows themselves. The preservation of cristobalite in the samples indicates that no significant post depositional metamorphism has affected these rocks.

Specimens cut from cores KH25-A and KH14-D were polished and imaged in reflected light with a Zeiss petrographic microscope and using backscattered scanning electron microscopy (BSEM) in the MIT Electron Microprobe Facility (see Supplementary Figure S3). In KH14-D, the dominant opaque phase is up to 300  $\mu\text{m}$  grain-size, cubic, pale-yellow colored and isotropic with a high reflectance and high polishing hardness, suggesting that it is pyrite ( $\text{FeS}_2$ ) and possibly other iron sulfides. In sample KH25-A, the

dominant opaque phase is mostly  $< 5 \mu\text{m}$  grain-size and appears grey in reflected light, has low reflectivity, and usually has an irregular grain shape, suggesting that it is a secondary alteration product. Wavelength-dispersive x-ray spectroscopy (WDS) showed that these grains are FeOOH (either goethite or lepidocrocite). While not observed with electron or optical microscopy, the peak unblocking temperatures of our specimens strongly suggest that magnetite ( $\text{Fe}_3\text{O}_4$ ) is present and is the dominant magnetization carrier. The lack of detection of magnetite using microscopy is expected given that hysteresis experiments (see Supplementary Figure S5) suggest that the magnetite grains present are single domain (SD) and therefore very small ( $< 500 \text{ nm}$ ) and so would be difficult to detect given the maximum resolution of the BSEM images ( $100 - 200 \text{ nm}$ ). In addition, the magnetic moment to volume ratio of our specimens ( $\sim 5 \times 10^{-8} \text{ Am}^2$  and  $\sim 1 \times 10^{-7} \text{ Am}^2$ , for KH14-D and KH25-A respectively) suggests that magnetite grains are present in very low concentration ( $< 1 \text{ ppb}$ ) and so would easily evade detection given the limited field of view of the Electron Microprobe.

### **Rock magnetism:**

The domain state of magnetite in a representative subset of bedded volcanic samples (KH01B, KH03F, KH14A and KH25A) was characterized by acquiring hysteresis curves. Hysteresis curves were measured in the Ross Laboratory, MIT Department of Materials Sciences and Engineering on an ADE model 1660 vibrating sample magnetometer (VSM). Hysteresis loops were acquired for applied fields of  $-1$  to  $1 \text{ T}$  and were corrected by removing any paramagnetic or diamagnetic signals (see Supplementary Figure S5). The corrected curves were used to ascertain the saturation magnetization ( $M_s$ ), the saturation remanent magnetization ( $M_{rs}$ ), the coercivity of remanence ( $H_{cr}$ ) and the coercivity ( $H_c$ ) for each sample (Supplementary Table S2). The ratio of  $M_{rs}/M_s$  and  $H_{cr}/H_c$  were calculated to assess the domain state of magnetite on a Dunlop-Day plot (1, 2) (see Supplementary Figure S6).

All of the samples plot in the pseudo-single domain (PSD) range on the Day plot suggesting they have reliable paleomagnetic recording properties. Samples KH01B and KH03F show ‘wasp-waisted’ hysteresis behavior, indicative of a mixture of single-domain (SD) and PSD carriers (3). Samples KH14A

and KH25A are approaching the sensitivity limit of the VSM but their hysteresis parameters also indicate a PSD state. These observations are consistent with the position of each sample on the Day plot; KH01B and KH03F have higher  $M_{rs}/M_s$  ratios and lower  $H_{cr}/H_c$  ratios placing them closer to the SD region.

### **Paleomagnetism:**

Most specimens contained a low temperature and low coercivity overprint (LTC) that was removed in AF steps and thermal steps below 200 °C (see Supplementary Figure S7). These components align with the present-day magnetic field direction in the study location and are interpreted to represent a recent overprint, either due to viscous remanent magnetization (VRM) (4), or chemical remanent magnetization (CRM) carried by goethite (see Supplementary Figure S8). In most of the specimens, stable, origin-trending, high temperature magnetization components with blocking temperatures consistent with magnetite were isolated in the thermal demagnetization steps between 500 °C and 580 °C (HT1, see Supplementary Figure S7 and Dataset S1). In a small subset of the specimens, stable, high-temperature components with blocking temperatures consistent with hematite were also isolated in the thermal steps between 600 °C and 680 °C (HT2, see Supplementary Figure S7 and Dataset S1). Samples from one site (KH06) had stable, origin-trending, magnetization components that were isolated in the thermal demagnetization steps between 200 °C and 350 °C that could be carried by titanomagnetite, pyrrhotite or goethite (LT2, see Supplementary Figure S7 and Dataset S1).

To test if the high temperature magnetic components (HT1 and HT2) represent a later overprinting magnetization, we conducted two conglomerate tests (5) using 46 clast specimens from an intraformational conglomerate unit at site KH12 (near the top of the section) and 23 clasts from another intraformational conglomerate unit at site KP3 (near the bottom of the section) (see Supplementary Figure S9). To test the randomness of clast magnetization directions, we performed a Watson (6) test on the distribution of tilt-corrected HT1 and HT2 components from the clast specimens. In the Watson (6) test, for the clast magnetization directions to be considered randomly distributed, the magnitude of the unit-vector sum ( $R$ ) of the clast magnetizations should be less than a critical value  $R_0$  which is defined as a function of the



number of clasts ( $n$ ). The distribution of tilt-corrected HT1 magnetization directions from  $n = 46$  igneous clasts at site KH12 yielded  $R = 6.305$ , which is considerably less than the 95% confidence critical value  $R_0 = 10.947$ . Likewise, the same test for the HT2 directions ( $n = 30$ ) from the clasts at site KH12 yields  $R = 7.572$ , less than the 95% confidence critical value  $R_0 = 8.840$ . The distribution of tilt-corrected HT1 components from igneous clasts at site KP3 ( $n = 18$ ) yielded  $R = 5.359$ , which is less than the 95% critical value  $R_0 = 6.848$ . These results demonstrate that to 95% confidence we cannot distinguish the distributions of the HT1 and HT2 magnetization directions among clasts in these conglomerates from a random distribution. This means that the conglomerate test passes, and both HT1 and HT2 magnetizations are not regional nor local overprints and therefore can be considered primary remanent magnetizations throughout our studied section.

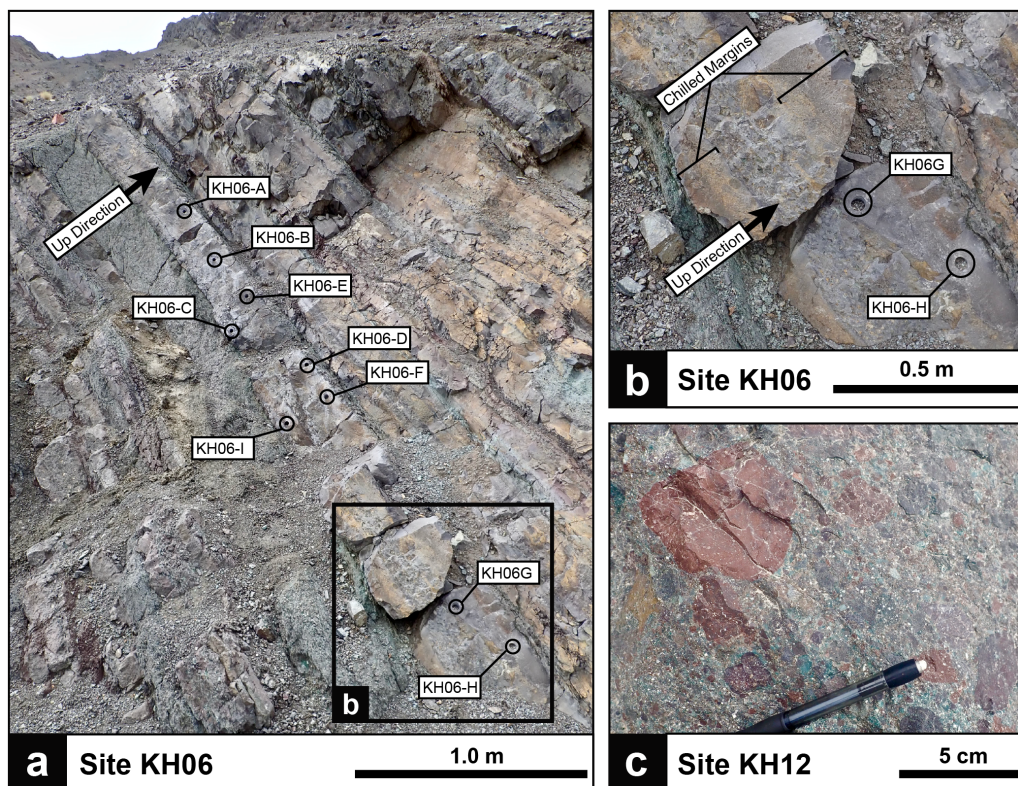
We interpret the HT1 component to represent a primary thermoremanent magnetization (TRM) acquired as the volcanic flows cooled below the magnetite Curie temperature. It is plausible that HT2 components either represent a primary TRM that was acquired as flows containing hematite cooled through 680 °C, or a chemical remnant magnetization (CRM) acquired during deuteric alteration converting magnetite to hematite. The most likely explanation for the contemporaneous formation of a magnetite TRM and hematite CRM in the Khardung volcanics is deuteric auto-metamorphism during dewatering and degassing of the water-rich felsic flows as they cooled. In either case, TRM or CRM, the fact that the HT2 magnetization directions pass the conglomerate test, and are almost exclusively near-parallel to HT1 directions from the same samples specimens, suggests that even if they are a CRM they were acquired very soon after the flows were formed and can still be considered primary. Primary site-mean directions were calculated from the HT magnetization directions that dominated the NRM using Fisher statistics (7) and corrected for bedding tilt. Three sites were excluded from further analysis due to their large site-mean uncertainties ( $\alpha_{95} > 15^\circ$ ) (8).

A bootstrap reversal test demonstrated that the two populations of site-mean directions are statistically antipodal to each other and are therefore highly likely to be the result of reversals in polarity of Earth's geodynamo (9) (see Supplementary Figure S10). The antipodal populations of site-mean directions

correlate to all but one of the expected geomagnetic polarity chrons (see main text Figure 3). It is unsurprising that the C27n chron is not observed because it is relatively short in duration ( $< 250$  kyr). Either there were no eruptions in that time or that we did not sample the flows in which it is preserved. Additionally, since our oldest age constraint is a maximum deposition age rather than an absolute deposition age, we conservatively assume that the two lowermost paleomagnetic sites are part of the C29n chron despite the age placing them within the opposite polarity C29r chron. It is also possible that these two sites actually correlate to C30n and C29r is not observed. If this is the case, it would only increase the number of magnetic polarity chrons preserved, strengthening our reversal test and increasing the time range included in our dataset. The passed reversal and conglomerate tests demonstrate that the section has not been overprinted by subsequent events that would have obliterated the reversal record (5).

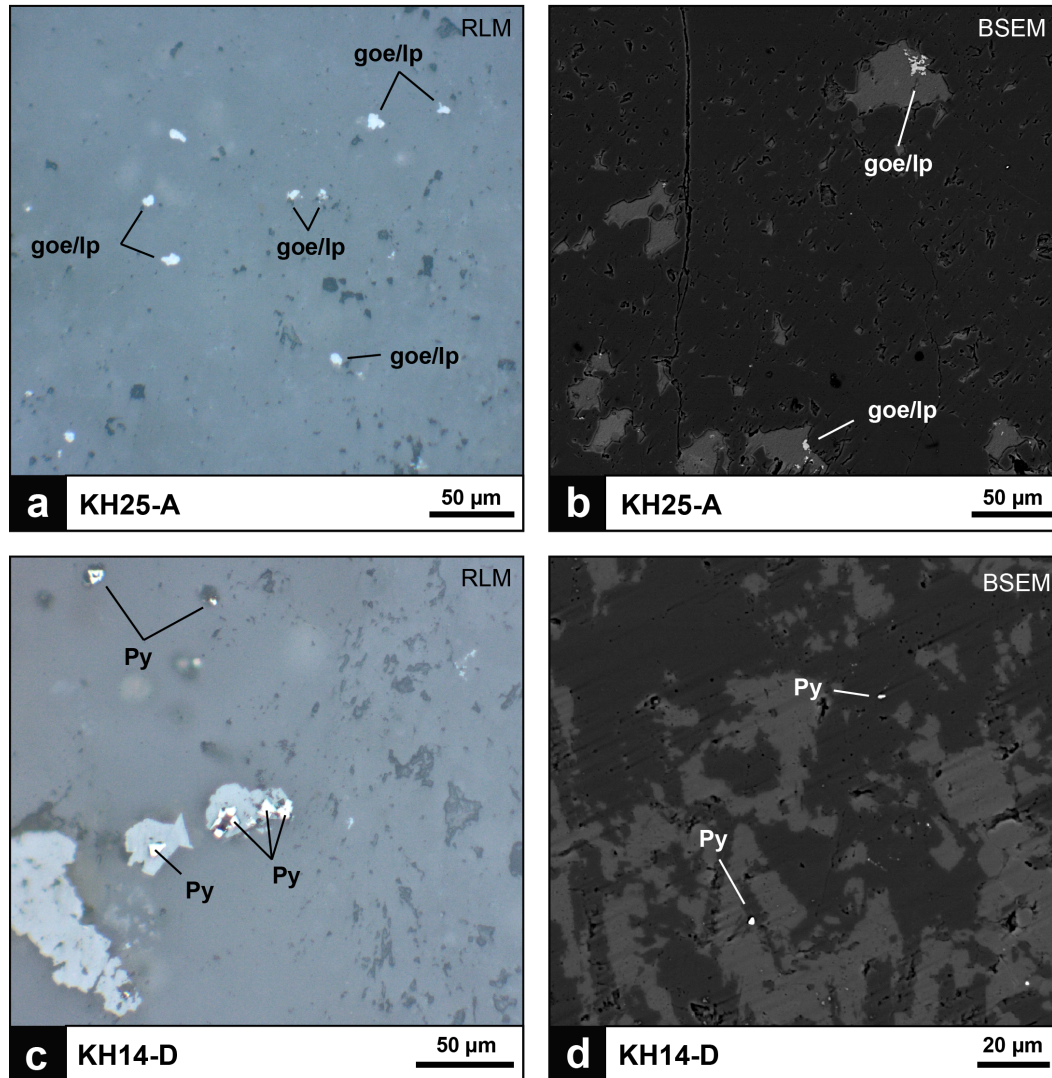
The time-averaged paleomagnetic pole was determined using northern hemisphere virtual geomagnetic poles because they are more likely to have a Fisher distribution than site-mean paleomagnetic directions (10). We used the quantile-quantile method (11) to assess whether the distribution of VGPs is indeed Fisher-distributed (See Supplementary Figure S11). We obtain the test statistics  $M_u = -0.904$  and  $M_e = 0.533$  that are less than their respective critical values [ $M_{u(critical)} = 1.207$ ,  $M_{e(critical)} = 1.095$  for  $n = 18$ ], indicating that the Khardung VGP distribution is consistent with a Fisher model (see Supplementary Figure S11). Northern hemisphere VGP positions for each site and the paleomagnetic mean pole for the Khardung volcanics are shown in Supplementary Table S3 and Supplementary Figure S12.



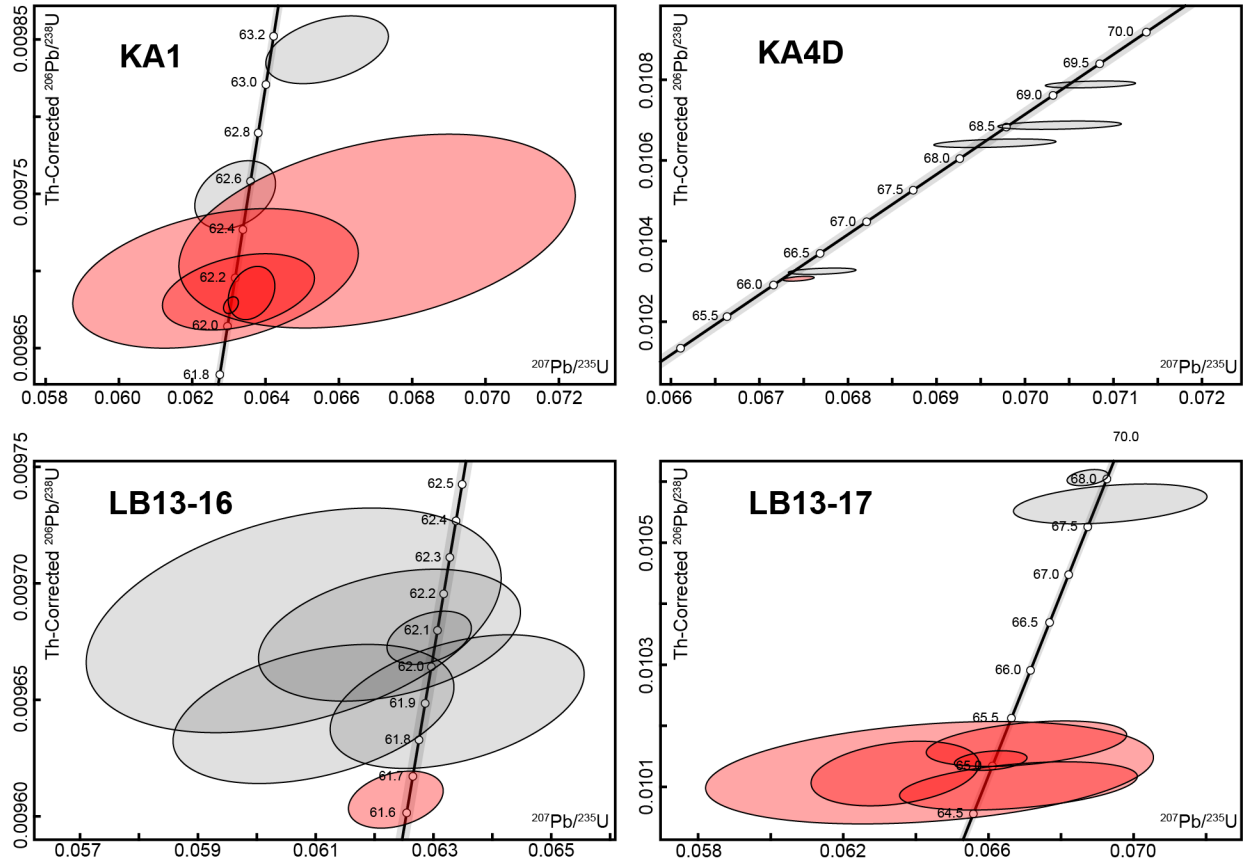


**Figure S2:** **a)** Lava flow interbedded with tuff and ash at site KH06 with locations of drill cores highlighted and up direction indicated. **b)** Blow up of lava flow sampled at site KH06 showing chilled lower and upper surfaces of the flow. **c)** conglomerate clasts that were drilled at site KH12 for a conglomerate test, transported clasts are exclusively rhyolitic and andesitic volcanic lithologies similar to other flows in the Khardung section.

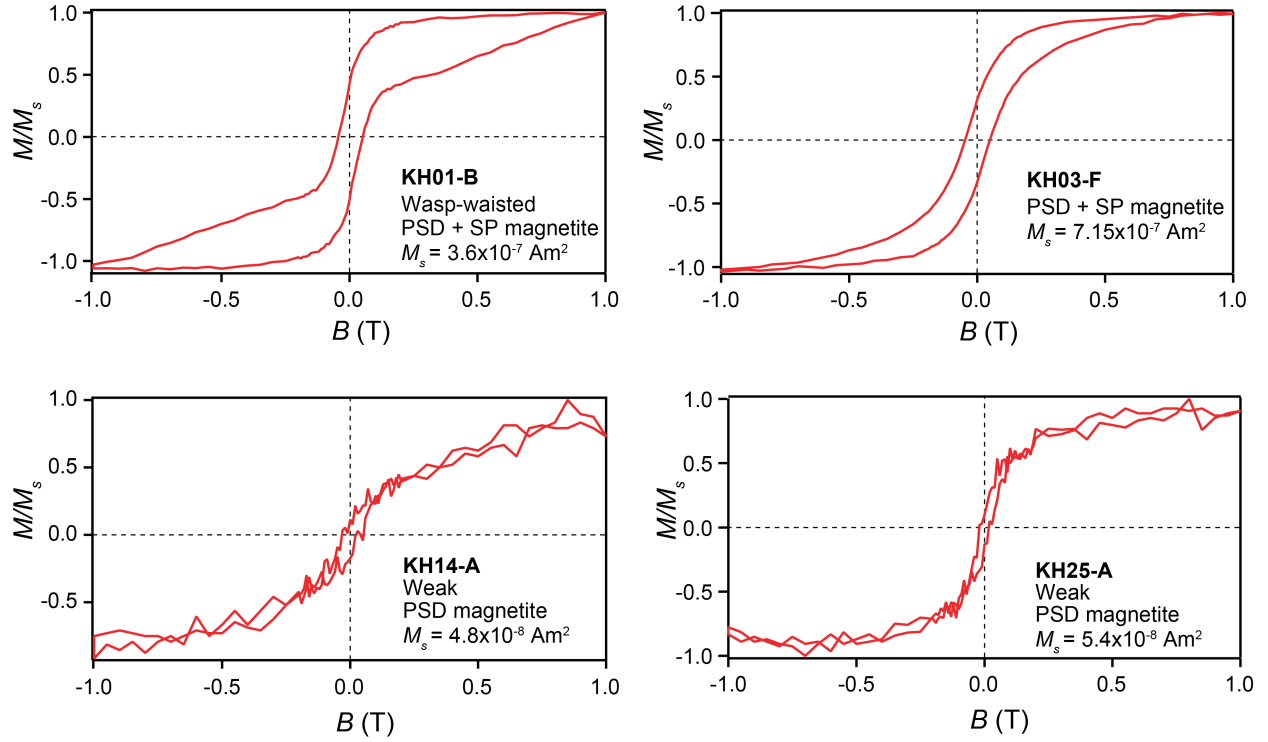




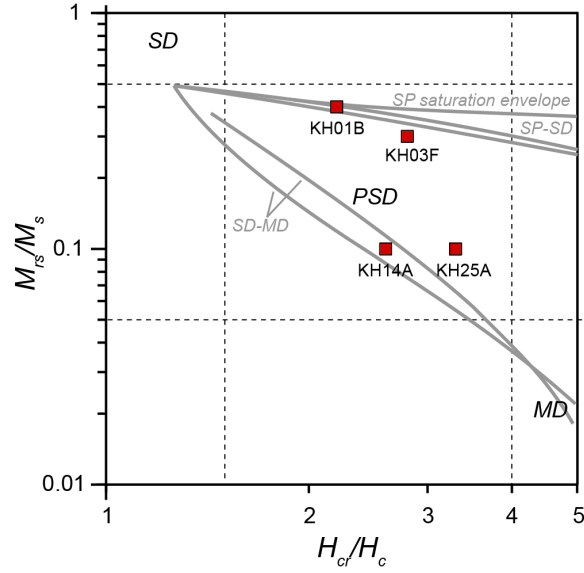
**Figure S3:** **a)** Reflected light microscope image of core sample KH25-A with goethite/lepidocrocite grains indicated. **b)** Backscattered scanning electron microscope (BSEM) image of KH25-A showing irregular goethite/lepidocrocite grains. **c)** Reflected light microscope image of core sample KH14-D showing pyrite. **d)** BSEM image of KH14-D.



**Figure S4:** U-Pb concordia plots and weighted mean Th-corrected  $^{206}\text{Pb}/^{238}\text{U}$  dates for KA1, KA4D, LB13-16 and LB13-17. Zircon grains contributing to the reported ages for each sample are shown in red. The dates from all of these samples are interpreted to represent the age of eruption/deposition.

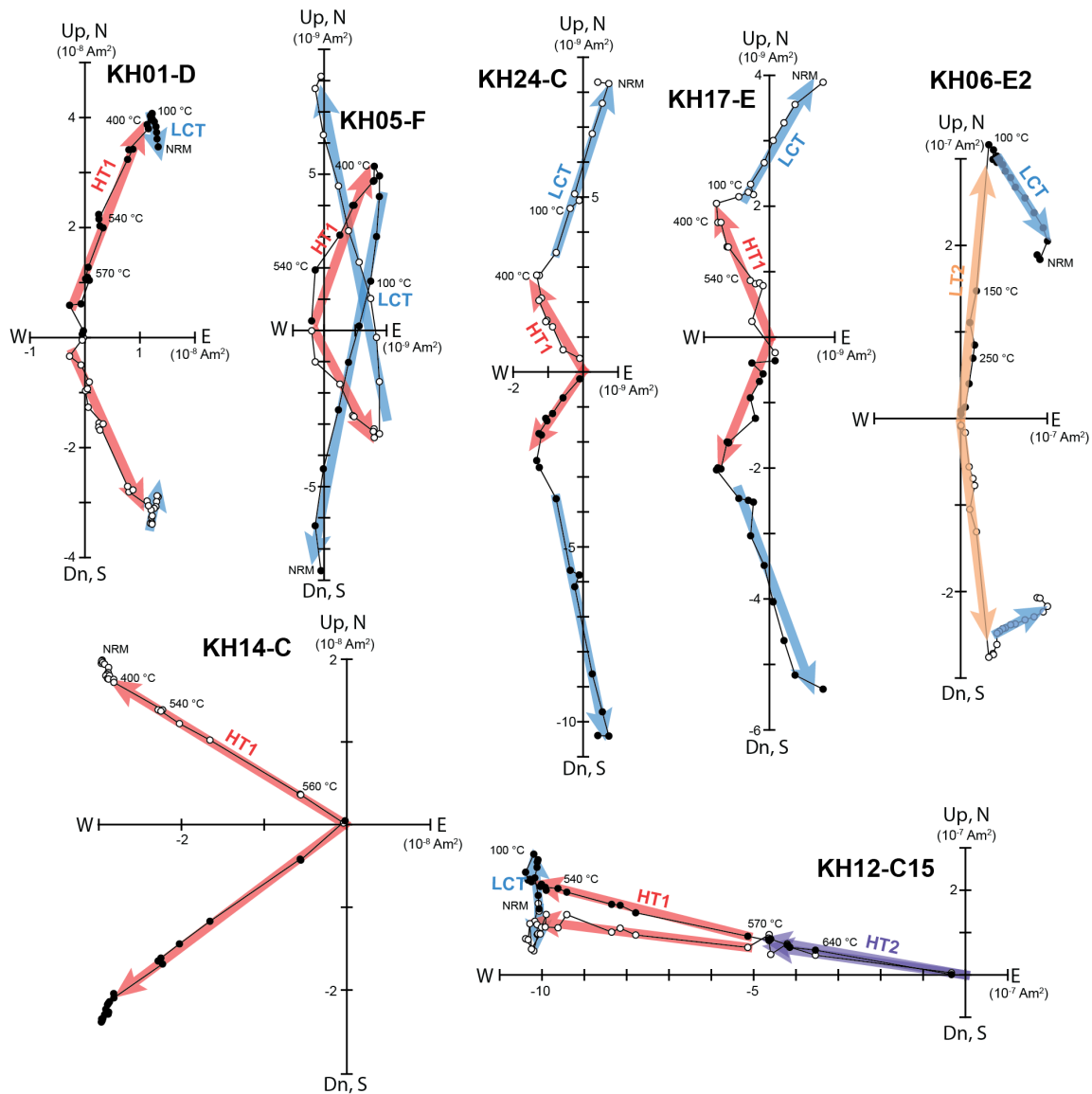


**Figure S5:** Hysteresis curves for the representative rhyolite samples KH01-B, KH03-F, KH14-A and KH25-D. Hysteresis loops were acquired for applied fields of  $B = -1.0 - 1.0$  T and were corrected by removing any paramagnetic or diamagnetic signals. The corrected curves were used to ascertain the saturation magnetization ( $M_s$ ), the saturation remanent magnetization ( $M_{rs}$ ), the coercivity of remanence ( $H_{cr}$ ) and the coercivity ( $H_c$ ) for each sample.

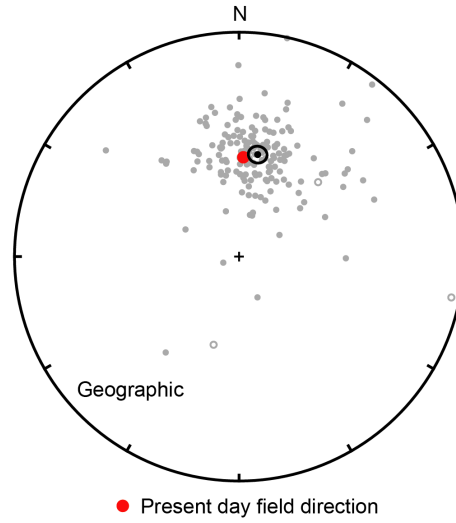


**Figure S6:** Day plot showing the ratio of saturation remanent magnetization ( $M_{rs}$ ) over saturation magnetization ( $M_s$ ), and the ratio of coercivity of remanence ( $H_{cr}$ ) and coercive force ( $H_c$ ) determined from hysteresis data for the representative samples KH01-B, KH03-F, KH14-A and KH25-D (see Supplementary Table S2). Regions are denoted to show the different domain states magnetite: MD = multidomain, PSD = pseudo-single domain, SD = Single domain. All of our samples are in pseudo-single domain field indicating that they are reliable paleomagnetic recorders (1, 2).

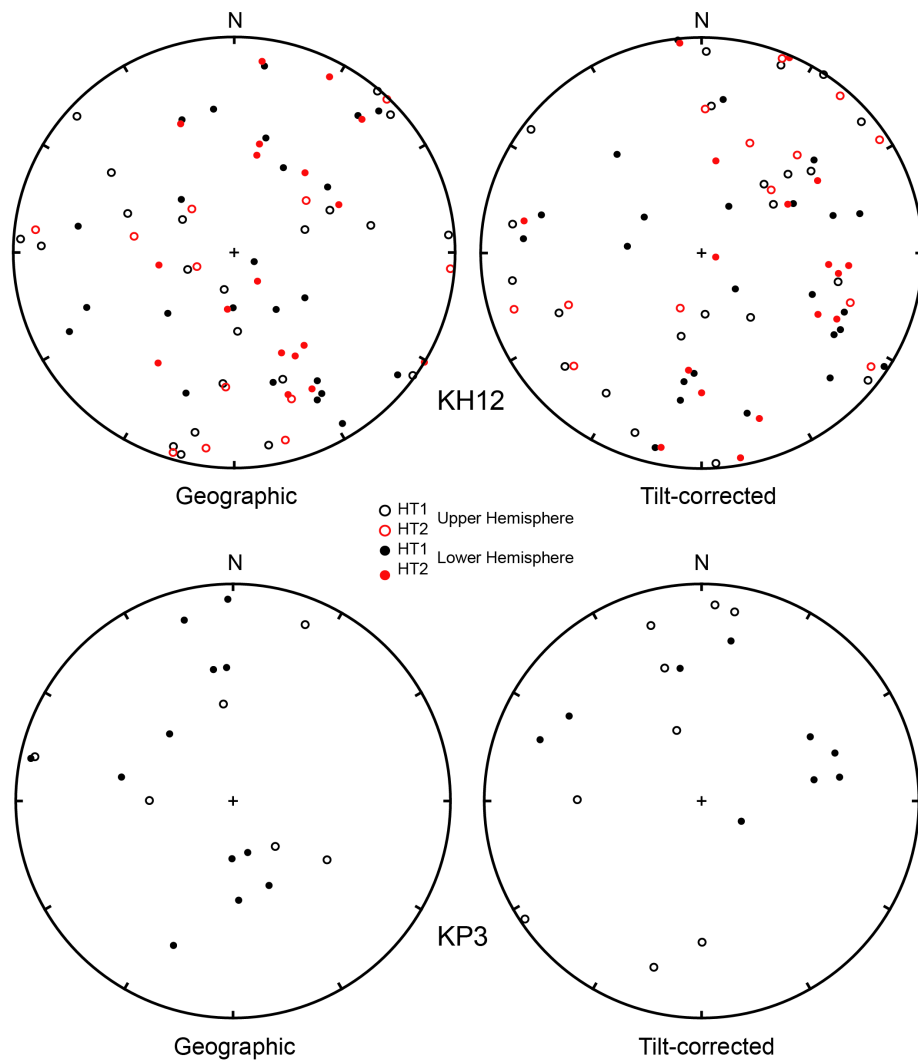




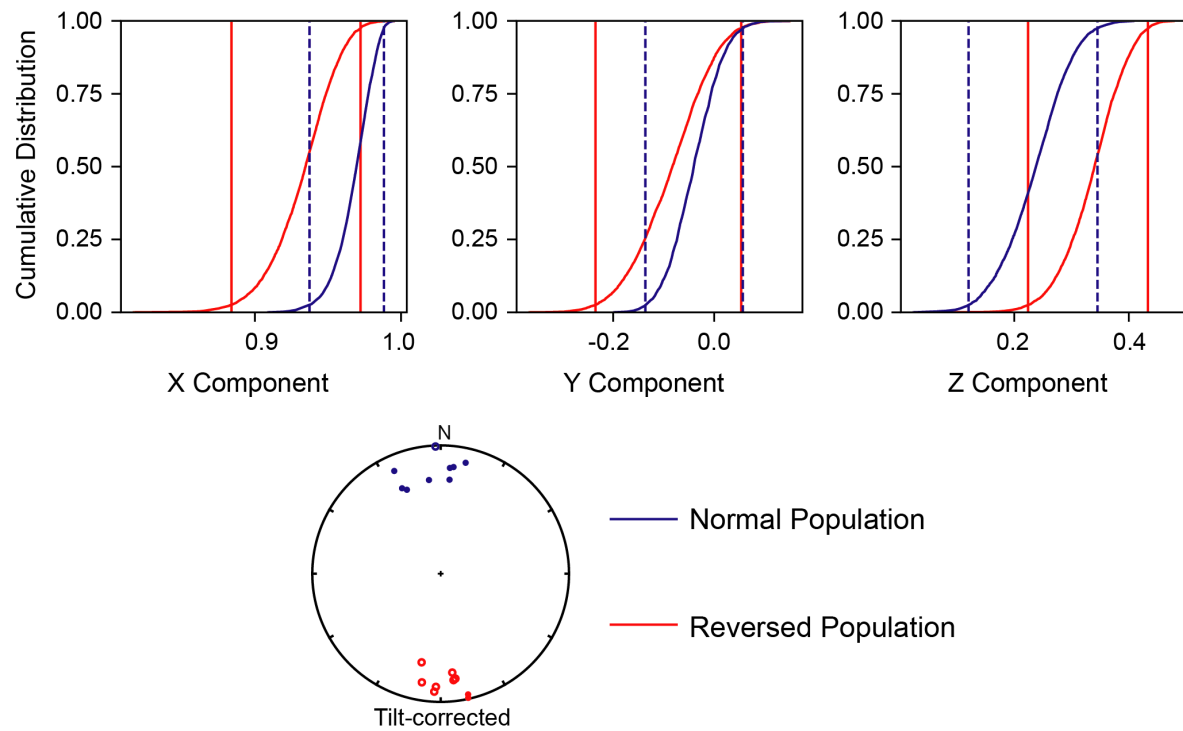
**Figure S7:** Zijderveld diagrams showing AF and thermal demagnetization of a representative set of specimens. Data is presented in geographic coordinates. Closed symbols represent north-south-east-west projections and open symbols represent up-down-east-west projections.



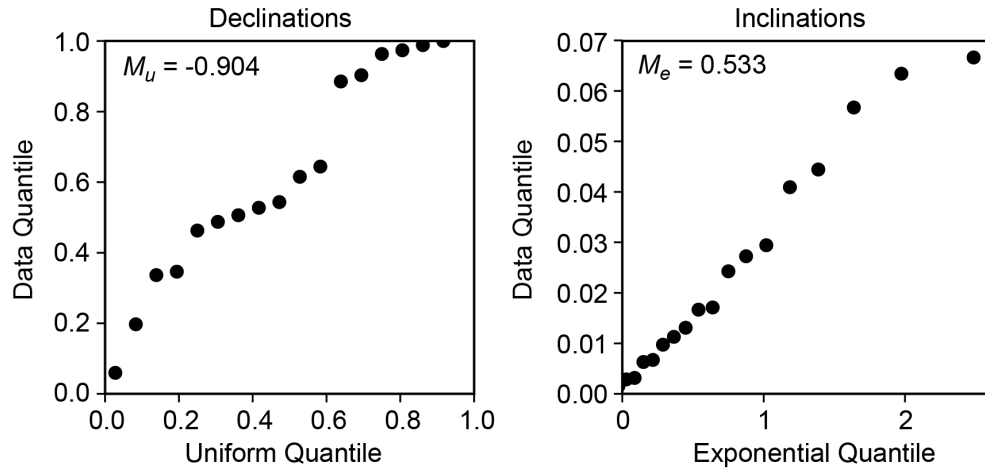
**Figure S8:** Equal area projection showing the LCT directions. The mean direction of the LCT components is almost equivalent to the modern-day field direction. Negative inclinations are projected onto the upper hemisphere and shown with open symbols, filled symbols represent positive inclinations.



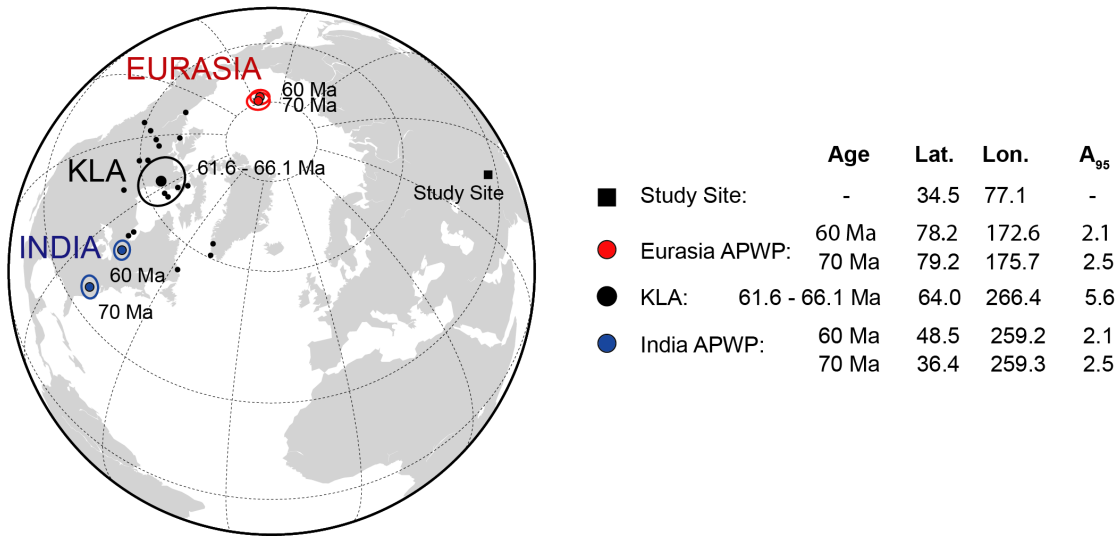
**Figure S9:** Equal area projection showing the HT1 and HT2 conglomerate test results. The HT1 directions of conglomerate clasts are shown in black and HT2 directions are shown in red. Negative inclinations are projected onto the upper hemisphere and shown with open symbols, while filled symbols represent positive inclinations projected onto the lower hemisphere. Both HT1 and HT2 directions are random with  $> 95\%$  certainty.



**Figure S10:** Cumulative distributions of Cartesian components of bootstrapped means from 500 pseudo-samples from tilt-corrected site-mean data (12). The 95% confidence intervals for the two populations overlap in X, Y, and Z. The two populations can be considered antipodal to one another because they share a common mean.



**Figure S11:** Quantile-quantile plots of longitudes and latitudes (in data coordinates) plotted against an assumed uniform and exponential distribution respectively. The data are Fisher distributed because the statistical parameters  $M_u$  and  $M_e$  are both less than the critical values.



**Figure S12:** Northern hemisphere virtual geomagnetic poles (VGPs) and mean paleomagnetic pole for the KLA compared to contemporaneous apparent polar wander path (APWP) poles for India and Eurasia (13). The poles do not coincide so the KLA was not moving rigidly with either India or Eurasia at 61.6 – 66.1 Ma.



**Table S1: CA-IDTIMS U-Pb Zircon Geochronology Results**

Frac.	Dates							Composition			Isotopic Ratios								
	<sup>206</sup> Pb/ <sup>238</sup> U*	2σ abs.	<sup>207</sup> Pb/ <sup>235</sup> U†	2σ abs.	<sup>207</sup> Pb/ <sup>206</sup> Pb†	2σ abs.	% Disc.§	Corr. Coef.	Th/ U#	Pb <sub>c</sub> **	Pb*/ Pb <sub>c</sub> ††	<sup>206</sup> Pb/ <sup>204</sup> Pb§§	<sup>208</sup> Pb/ <sup>206</sup> Pb##	<sup>206</sup> Pb/ <sup>238</sup> U*##	2σ %	<sup>207</sup> Pb/ <sup>235</sup> U##	2σ %	<sup>207</sup> Pb/ <sup>206</sup> Pb*##	2σ %
<b>KA4D</b>																			
z1	66.088	0.039	66.26	0.17	75.7	5.6	12.78	0.475	0.63	0.23	102.2	5893	0.202	0.0102915	0.058	0.06744	0.26	0.04755	0.23
z4	89.86	0.061	90.17	0.47	101	12	10.74	0.366	0.66	0.37	44.9	2581	0.212	0.0140242	0.067	0.09287	0.55	0.04805	0.52
z5	68.51	0.067	69.05	0.65	90	23	24.36	0.297	0.85	0.2	24	1327	0.271	0.010672	0.097	0.07037	0.98	0.04784	0.96
z6	66.20	0.051	66.51	0.36	80	13	17.86	0.378	0.7	0.34	41.7	2377	0.223	0.01031	0.077	0.0677	0.56	0.04764	0.53
z7	68.22	0.068	68.36	0.65	76	22	9.97	0.393	0.83	0.22	25.2	1396	0.265	0.010627	0.1	0.06964	0.98	0.04755	0.95
z8	69.14	0.057	69.39	0.48	80	16	13.66	0.34	1.18	0.28	37.2	1891	0.376	0.0107733	0.079	0.07072	0.71	0.04763	0.69
<b>LB13-17</b>																			
z1	64.93	0.34	62.5	2.2	-27	85	343.56	0.293	0.81	0.43	7.7	442	0.259	0.010111	0.52	0.0635	3.6	0.0456	3.5
z3	68.015	0.081	67.56	0.53	54	19	-26.57	0.319	1.06	0.32	34.5	1803	0.34	0.010596	0.12	0.0688	0.81	0.04711	0.78
z5	64.93	0.53	63.4	5.8	9	220	-601.07	0.295	1.14	0.65	3	172	0.365	0.010113	0.83	0.0645	9.5	0.0462	9.3
z7	67.75	0.21	68.2	2.5	85	88	20.46	0.358	0.87	0.44	6.3	360	0.279	0.010553	0.3	0.0694	3.8	0.0477	3.7
z8	65.062	0.099	65.01	0.94	66	35	1.35	0.273	0.95	0.37	16.6	902	0.304	0.010133	0.15	0.06612	1.5	0.04735	1.5
z10	64.79	0.25	65.8	3.1	103	110	37.37	0.391	1.07	0.43	5.2	289	0.343	0.010091	0.39	0.0669	4.8	0.0481	4.7
z11	65.24	0.23	66	2.6	95	93	31.36	0.382	0.92	0.28	6.2	348	0.295	0.01016	0.36	0.0671	4.1	0.0479	3.9
<b>KA1</b>																			
z1	62.17	0.29	61.6	3.7	41	140	-51.54	0.357	1.04	0.42	4.1	234	0.333	0.009679	0.47	0.0625	6.2	0.0469	6
z2	62.36	0.4	65.7	5.1	192	180	67.64	0.362	1.08	0.49	3	174	0.346	0.00971	0.64	0.0669	8	0.05	7.8
z3	62.075	0.035	62.08	0.19	65.6	7.3	5.53	0.341	0.64	0.19	76.6	4420	0.204	0.0096629	0.056	0.06305	0.32	0.04735	0.3
z4	62.52	0.14	62.2	1	51	40	-21.53	0.323	1.02	0.34	15.7	841	0.326	0.009734	0.23	0.0631	1.7	0.04706	1.7
z5	62.13	0.11	62.61	0.6	84	23	26.42	0.26	0.63	0.3	25.9	1507	0.203	0.009671	0.18	0.0636	0.99	0.04772	0.96
z7	63.13	0.14	64.5	1.6	120	58	47.59	0.393	0.72	0.57	9.3	540	0.229	0.009828	0.22	0.0656	2.5	0.0485	2.4
z8	62.13	0.16	62.2	2	69	75	10.01	0.354	0.59	0.5	6.8	416	0.188	0.009671	0.25	0.0632	3.2	0.0474	3.2
<b>LB13-16</b>																			
z3	62.14	0.31	59.8	3.3	-28	140	318.91	0.372	0.74	0.42	4.4	264	0.237	0.009674	0.49	0.0607	5.7	0.0455	5.6
z4	61.89	0.19	60.2	2.3	-5	91	1316.72	0.383	0.53	0.37	6.1	379	0.168	0.009633	0.31	0.061	3.9	0.046	3.8
z5	62.09	0.18	61.4	2.1	37	82	-69.63	0.343	0.96	0.58	7	392	0.307	0.009667	0.29	0.0623	3.5	0.0468	3.4
z7	61.641	0.074	61.48	0.76	58	30	-6.03	0.296	0.91	0.28	20.4	1112	0.292	0.009596	0.12	0.06242	1.3	0.0472	1.2
z8	62.092	0.069	62.02	0.67	62	26	0.69	0.347	0.72	0.33	22.1	1261	0.231	0.009666	0.11	0.06299	1.1	0.04728	1.1
z10	61.93	0.18	62.5	2	87	77	28.59	0.412	0.7	0.54	7.3	429	0.224	0.009641	0.29	0.0635	3.4	0.0478	3.2

\* Corrected for initial Th/U disequilibrium using radiogenic <sup>208</sup>Pb and Th/U<sub>[Magma]</sub>=2.8±0.5.

† Isotopic dates calculated using the decay constants λ<sub>238</sub> = 1.55125E-10 and λ<sub>235</sub> = 9.8485E-10 (14).

§ % discordance = 100 - (100 \* (<sup>206</sup>Pb/<sup>238</sup>U date) / (<sup>207</sup>Pb/<sup>206</sup>Pb date))

# Th contents calculated from radiogenic <sup>208</sup>Pb and the <sup>207</sup>Pb/<sup>206</sup>Pb date of the sample, assuming concordance between U-Th and Pb systems.

\*\* Total mass of common Pb.

†† Ratio of radiogenic Pb (including <sup>208</sup>Pb) to common Pb.

§§ Measured ratio corrected for fractionation and spike contribution only.

## Measured ratios corrected for fractionation, tracer and blank.

**Table S2: Hysteresis parameters calculated for a representative suite of samples.**

Parameter	KH01-B	KH03-F	KH14-A	KH25-A
$M_s$ (Am <sup>2</sup> )	$1.51 \times 10^{-7}$	$2.29 \times 10^{-7}$	$6.72 \times 10^{-9}$	$5.40 \times 10^{-9}$
$M_{rs}$ (Am <sup>2</sup> )	$3.60 \times 10^{-7}$	$7.15 \times 10^{-7}$	$4.80 \times 10^{-8}$	$5.40 \times 10^{-8}$
$H_{cr}$ (mT)	100	140	90	65
$H_c$ (mT)	45	50	35	20
$M_{rs}/M_s$	0.4	0.3	0.1	0.1
$H_{cr}/H_c$	2.2	2.8	2.6	3.3
$M_s$ = saturation magnetization				
$M_{rs}$ = saturation remanent magnetization				
$H_{cr}$ = the coercivity of remanence				
$H_c$ = coercivity				

**Table S3: Site Mean Directions and Virtual Geomagnetic Pole Positions**

Site	Slat (°)	Slon (°)	Bedding Strike (°) /Dip (°)	Comp.	$n/n_0$	$D_g$ (°)	$I_g$ (°)	$D_s$ (°)	$I_s$ (°)	$\alpha_{95}$ (°)	$K$	$R$	$Plat$ (°)	$Plon$ (°)	$dm$ (°)	$dp$ (°)	Pol.
KH01	34.45430	77.72338	308.9/53.8	LCT	4/8	18.7	51	27.2	-0.6	17	30	3.9007					
				HT1	8/8	155.1	-48.1	182.4	-12.7	2.3	565	7.9876	-61.9	72.7	2.3	1.2	R
				HT2	5/8	199.4	-56.1	209.2	-4.2	10.2	57	4.9295					
KH02	34.45970	77.71894	293.6/50.5	LCT	6/6	15.8	71.9	22.5	21.6	60.5	2	3.7057					
				HT1	6/6	143.9	-47.1	167.8	-12.6	8.3	67	5.9253					
				HT2	5/6	155.3	-36.6	167.5	0.9	4.8	252	4.9841	-53.2	98.9	4.8	2.4	R
KH03	34.45314	77.72160	310/49.4	LCT	7/8	3.2	48.7	17.5	5.7	15.2	17	6.6403					
				HT1	7/8	146.1	-42.1	173.1	-17.5	10.2	36	6.8327	-63.7	93.3	11	5.5	R
				HT2	4/8	139.4	-52.6	179.4	-27.3	19.8	23	3.8669					
KH04	34.45097	77.71850	312.8/46.8	LCT	8/8	11.1	55	25.9	12.4	23	7	6.9609					
				HT1	6/8	135	-33.8	161	-22.4	20.7	11	5.5635					
				HT2	7/8	142.9	-43.3	173.3	-23.4	6.8	81	6.9256	-66.9	94.6	7.2	3.9	R
KH05	34.45915	77.71881	298/51.1	LCT	12/12	354.7	48.8	7.7	3.2	6.4	47	11.7681					
				HT1	12/12	140.8	-51.2	172	-18.5	4.4	98	11.8882	-64	96	4.6	2.4	R
KH06	34.45303	77.72109	309.1/45.2	LCT	11/12	353.2	60.7	18.1	22.5	16.7	8	9.8158					
				LT2	10/12	169.4	-50.4	189.8	-15.1	2.7	321	9.9720	-61.8	56.8	2.8	1.4	R
KH07	34.44875	77.71827	313.7/59	LCT	5/5	9.8	56.4	26.8	2.6	13.1	35	4.8895					
				HT1	5/5	297.6	32.6	338.1	30.2	14.3	30	4.8657	63.1	310.5	16	8.8	N
KH11	34.44809	77.71645	321.3/62.1	LCT	5/5	31.1	45.6	34.4	-20.7	23.8	11	4.6453					
				HT1	5/5	152.1	-28.7	167.2	3.9	8.4	84	4.9525	-51.7	98.6	8.4	4.2	R
KH12	34.45733	77.72047	308/53.8	LCT	10/10	6.5	52.7	20.5	3.6	5.2	86	9.8950					
				HT1	10/10	160.2	-46.6	183.1	-8.7	7.4	44	9.7954	-59.8	71.6	7.5	3.8	R
KH14	34.44967	77.71896	300.8/46.3	HT1	6/6	302.5	41.9	335.7	27.8	7.4	83	5.9394	60.6	311.8	8.1	4.4	N
KH16	34.44802	77.71579	321.3/62.1	LCT	10/10	11.7	57.3	32.4	2.2	5.8	71	9.8737					
				HT1	10/10	324.8	44.7	5	18.4	6.3	60	9.8498	64.6	246.2	6.5	3.4	R
KH17	34.44800	77.71570	321.3/62.1	LCT	10/10	348.1	57.1	31.4	2.7	6.6	54	9.8339					
				HT1	10/10	327	46	6.9	17.4	3.1	236	9.9619	63.7	242.2	3.3	1.7	R
KH18	34.44807	77.71700	319.1/53.1	LCT	10/10	37.7	45.3	56.1	-6.9	26.9	4	7.8483					
				HT1	4/10	116	-23	146.6	-33.3	180	2	2.3306					
KH19	34.44817	77.71729	315/48.2	LCT	10/10	11.5	47.8	24	5	8.4	34	9.7325					
				HT1	8/10	317.5	43.8	352.9	27.3	6.6	71	7.9017	69	277.2	7.2	3.9	N
KH20	34.44834	77.71790	313.7/59	LCT	9/9	5.3	49.1	20.6	-2.1	6.8	59	8.8636					

				HT1	7/9	314.3	23.6	335.7	13.2	10.9	32	6.8111	54.3	302.2	11	5.7	N
KH21	34.44856	77.71812	313.7/59	LCT	10/10	15	45.9	24.7	-8.2	9.8	25	9.6445					
				HT1	7/10	339.1	26.8	350.6	-3.9	24.9	7	6.1234					
KH22	34.44797	77.71533	318.5/56.1	LCT	9/9	18.3	55.3	28.2	0.5	8.5	38	8.7879					
				HT1	5/9	317.8	50.8	5.4	27.3	9.8	63	4.9360	69.4	242.7	11	5.8	N
KH23	34.44799	77.71577	321.3/62.1	LCT	10/10	6.3	50.5	25.7	-1.2	8.1	36	9.7530					
				HT1	6/10	337.3	48.9	12.7	12.4	12.4	30	5.8339	59.5	232.2	13	6.4	N
KH24	34.44843	77.71773	311.6/67.2	LCT	10/10	359.9	52.2	18.5	-5.8	9.6	27	9.6605					
				HT1	8/10	335.3	39.5	357.7	-0.9	9.9	32	7.7819	55	261.7	9.9	5	N
KH25	34.46505	77.71215	301.5/50.5	LCT	10/13	17.3	49.2	23.1	-0.1	22.1	10	5.5076					
				HT1	13/13	145.1	-69.9	192.1	-30	2.5	270	12.9555	-68.7	44	2.8	1.5	R
KH26	34.44878	77.71796	313.7/59	LCT	10/10	4.2	53	22.3	1.5	11.5	19	9.5170					
				HT1	9/10	285	32.3	333.8	40.3	83.1	1	3.0624					
<b>N/N0 = 18/21</b>												<b>Mean Pole:</b>	<b>64.0</b>	<b>266.4</b>	<b>A<sub>95</sub> = 5.6</b>	<b>K = 38.5</b>	

Site = Identification number for site or cooling unit.

Site Lat./Lon. = GPS Latitude and longitude of sampling location.

Bedding Strike/Dip = mean bedding orientation at site used for tilt-correction.

Comp. = Component name, LCT: low temperature/coercivity, LT2: origin-trending low temperature, HT1: high temperature magnetite, HT2: high temperature hematite.

$n/n_0$  = ratio of number of specimen directions contributing to mean to number of specimens measured.

$D_g/D_s$  = Declination: geographic/stratigraphic coordinates.

$I_g/I_s$  = Inclination: geographic/stratigraphic coordinates.

$\alpha_{95}$  = 95% uncertainty radius around mean.

$K$  = Precision parameter.

$R$  = Magnitude of the vector sum of directions contributing to the mean.

$Plat/Plon$  = Latitude and longitude positions of virtual geomagnetic poles (VGPs) determined from the site mean directions.

$dm/dp$  = semi-axis of error ellipse surrounding VGP.

Pol. = polarity of VGP, N: normal, R: reversed.

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